

**M1.(a)** Star much brighter than reflected light from planet ✓

Or

Planet very small and distant – subtends very small angle compared to resolution of telescopes

1

(b) Planet and star orbit around common centre of mass that means the star moves towards/away from Earth as planet orbits ✓

1

Causes shift in wavelength of light received from star ✓

1

(c) Light curve showing constant value with dip ✓

1

When planet passes in front of star (as seen from Earth), some of the light from star is absorbed and therefore the amount of light reaching Earth reduced ✓

1

Apparent magnitude is a measure of the amount of light reaching Earth from the star ✓

1

[6]

**M2.** (a) (use of  $m - M = 5 \log(d/10)$  gives)  
 $3.54 - (-20.62) = 5 \log(d/10)$  (1)

$d = 6.7(9) \times 10^5 \text{pc}$  (1)

2

(b) use of  $\frac{\Delta\lambda}{\lambda} = -\frac{v}{c}$  (1)

$$\Delta\lambda = -\frac{0.21121 \times 105 \times 10^3}{3.0 \times 10^8} = -7(.4) \times 10^{-5}$$

$$\lambda' = 0.21121 - 7(.4) \times 10^{-5} = 0.21114m \text{ (1)}$$

(allow C.E. for incorrect value of  $\Delta\lambda$ )

2

(c)  $t\left(= \frac{d}{v}\right) = \frac{6.79 \times 10^5 \times 3.08 \times 10^{16}}{105 \times 10^3}$  (1)

$$= 2.0 \times 10^{17}s \text{ (1)}$$

$$(1.99 \times 10^{17}s)$$

(allow C.E. for value of  $d$  from (a))

2

[6]

**M3.** (a) (use of  $\frac{\Delta\lambda}{\lambda} = -\frac{v}{c}$  gives)  $\frac{(660.86 - 656.28)}{656.28} = (-)\frac{v}{3.0 \times 10^8}$  (1)

$$v = (-)2094 \text{ km s}^{-1} \text{ (1)}$$

2

- (b) graph to show:  
 correct plotting of points (1)  
 straight line through origin (1)

$$H = \frac{v}{d} = \text{gradient} = 70 \text{ km s}^{-1} \text{ Mpc}^{-1} \text{ (1)}$$

(must show evidence of use of graph in calculation)

3

[5]

- M4.** (a) (i) correct shape of graph (steeper on left of peak) **(1)**  
 (ii) region to left of peak **(1)**  
 (iii) ozone **(1)**  
 (iv) lower temperature, shifts peak ( $\lambda_{\text{max}}$ ) to longer wavelengths **(1)**  $\lambda_{\text{max}} T = \text{constant}$  **(1)**

max 4

- (b) (i) (use of  $f = \frac{c}{\lambda}$  gives)  $f \left( = \frac{3 \times 10^8}{2.7} \right) = 1.1 \times 10^8 \text{ Hz}$ ,  
 (in range) **(1)**  
 (ii) (double) Doppler **(1)**  
 (iii) (reflection off moving object gives double Doppler),  
 frequency shift = 150 Hz

$$v = \frac{150 \times 3 \times 10^8}{1.1 \times 10^8} \quad \mathbf{(1)}$$

(allow C.E. for shift = 300 Hz)

$$= 4.1 \times 10^2 \text{ m s}^{-1} \text{ (towards each other) } \mathbf{(1)}$$

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[9]

**M5.(a)** (i)  $\Delta\lambda = \frac{\lambda v}{c} \quad \mathbf{(1)}$

$$(ii) \quad \Delta\lambda = -\frac{\lambda v}{c} \quad (1)$$

(2)

$$(b) \quad (i) \quad \text{total difference in wavelength} = \frac{2\lambda v}{c} \quad (1)$$

$$v = \frac{7.8 \times 10^{-12} \times 3.0 \times 10^8}{589 \times 10^{-9} \times 2} = 1986 \text{ [or } 2.0 \times 10^3 \text{] m s}^{-1} \quad (1)$$

$$(ii) \quad \omega = \frac{v}{r} = \frac{1986}{7.0 \times 10^8} \quad (1)$$

$$= 2.8 \times 10^{-6} \text{ rad s}^{-1} \quad (1)$$

(4)

[6]