

A-Level Physics

Particles

Question Paper

Time available: 74 minutes Marks available: 52 marks

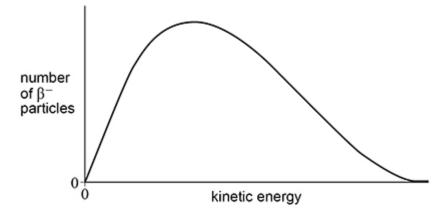
www.accesstuition.com

	1
ı	١.

Carbon-14 decays into nitrogen-14 with the release of a beta (β ⁻) particle and an antineutrino ($\nu_{\rm e}$).

(a) State the change of quark character in $\beta^{\scriptscriptstyle -}$ decay.

(b) The diagram below shows the distribution of kinetic energies of β^- particles from the decay of carbon-14.



Explain how the figure above supports the existence of the antineutrino.

(2)

The existence of the antineutrino was confirmed by experiments in which antineutrinos interact with protons. The equation for this interaction is:

$$\overline{v_e}$$
 + p \rightarrow e⁺ + X

(c) Identify particle X.

(1)

(d) The positron released in this interaction is annihilated when it encounters an electron. A pair of gamma photons is then produced.

Particle **X** can be absorbed by a nucleus. This produces another gamma ray.

The table below contains data for three gamma photons detected during an antineutrinoproton interaction experiment.

Gamma photon	Photon energy / J
G1	5.0×10^{-14}
G2	6.6×10^{-14}
G3	1.0×10^{-13}

Deduce which of the three gamma photons could have been produced by positron annihilation.

(3)

(Total 7 marks)

2. Two stable isotopes of helium are ${}_{2}^{4}$ He and ${}_{2}^{3}$ He.

(a) An atom of ${}^4_2\text{He}$ is produced in a rock that contains uranium. It is produced following the radioactive decay of a ${}^{238}_{92}\text{U}$ atom. The decay also creates an atom of thorium (Th).

Write an equation for the decay of $^{238}_{92}$ U.

$$^{238}_{92}U \rightarrow$$

(b) A ${}_{2}^{3}$ He nucleus can be produced by the decay of a tritium nucleus ${}_{1}^{3}$ H.

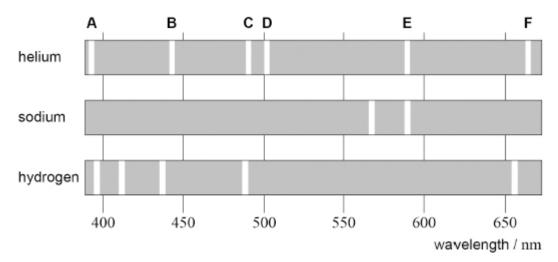
State and explain which exchange particle is responsible for this decay.



(2)

Helium was discovered by analysing the light in the absorption spectrum of the Sun.

The figure below shows the positions of the brightest lines, labelled **A** to **F**, in the **emission** spectrum of helium. The brightest lines in the emission spectra of sodium and hydrogen are also shown.



Before helium was identified, some scientists suggested that the lines of the helium spectrum seen in the absorption spectrum of the Sun were due to the presence of sodium and hydrogen.
Discuss, with reference to the lines ${\bf A}$ to ${\bf F}$ in the figure above, the evidence for and against this suggestion.
Calculate, in $eV,$ the change in energy level responsible for the spectral line labelled $\mbox{\bf E}$ in the diagram above.
change in energy level =eV

(3) (Total 12 marks)

A sigma-plus (Σ^+) particle and an unidentified particle **Y** are produced by the strong interaction between a positive pion (π^+) and a proton (p).

This interaction is represented by the equation:

$$\pi^+$$
 + p $\rightarrow \Sigma^+$ + **Y**

(a) Complete the table below to show the baryon number B, charge Q and strangeness S for the particles in this interaction.

	π+	p	Σ+	Y
В				0
Q	+1	+1	+1	
S				+1

		above has the quark structure uus?	
Tick	(√) one box.		
π+			
p			
Σ+			
Y			

(Total 6 marks)

•	(a)	Identify the number of neutrons in a nucleus of polonium-210 $\binom{210}{84}$ Po.	
		Tick (✓) one box.	
		84	
		126	
		210	
		294	
			(1)
	(b)	A polonium-210 nucleus is formed when a stationary nucleus of bismuth-210 decays. A beta-minus (β^-) particle is emitted in this decay.	
		Outline, with reference to $\beta^{\scriptscriptstyle -}$ decay, why bismuth-210 and polonium-210 have different proton numbers.	
			(2)

The kinetic energies of β^- particles emitted from a sample of bismuth-210 are analysed. These β^- particles have a range of kinetic energies.

The total energy released when each nucleus of bismuth-210 decays to a nucleus of polonium-210 is 1.2 MeV.

Figure 1 shows the variation with E_k of the number of β^- particles that have the kinetic energy E_k .

number of β^- with $E_{\mathbf{k}}$

 E_k / MeV

during β ⁻ dec	rigure i sup	oport the hy	pomesis m	at a triiru par	licie is produc

(c)

(d)) This third	particle is	an electron	antineutrino

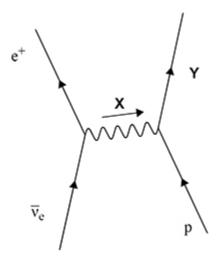
eta^- decay.	Explain why an electron antineutring	o, rather than ar	n electron neutrir	no, is produced	during
	β ⁻ decay.				

		·

(e) A large tank of water is used as part of an electron antineutrino detector. An electron antineutrino $\overline{\mathbf{v}}_{\mathbf{e}}$ enters the tank and interacts with a proton (p).

Figure 2 represents this interaction.

Figure 2



Identify X and Y.

(2)

ou shou	ld identify the lept	on in the molecule of water	ollides with this lepton. In your answe
		on in the molecule of water.	
he rang	e of the electroma	gnetic interaction is infinite.	
	below gives the rateraction.	ange of the strong nuclear	interaction and the range of the wea
	noraction.		1
lı	nteraction	Range / m	
str	ong nuclear	10 ⁻¹⁵	
W	eak nuclear	10 ⁻¹⁸	
		on or the electron antineutri er before interacting.	no is likely to travel the shorter
istance	in the tank of wate	er before interacting.	
			

E	
5.	

Cosmic rays are high-energy particles that come from space. Most of these particles are protons. There are other particles in cosmic rays, including atomic nuclei.

The table below gives the data for one particular nucleus **X**.

Mass / kg	8.02 × 10 ⁻²⁶
Specific charge / C kg ⁻¹	4.39 × 10 ⁷
Kinetic energy / MeV	215

(a) Determine the number of neutrons in nucleus X.

(3)

(b) Calculate the speed of **X**. Ignore relativistic effects.

(3)

A pion (π^+) and a kaon (K^+) are produced when cosmic rays interact with the upper atmosphere.

The K ⁺ decays to produ	ice an anti-muon and a muon neutrino.	
Explain how strangene	ss applies in this decay.	